# Modelling an extra dimension with domain-wall branes

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#### Overview

Physics beyond the standard model: extra dimensions.

- Brane  $\leftrightarrow$  domain wall (topological defect).
- Gravity & SM particles trapped to brane.
- At low energies,  $5D \rightarrow 4D$ .

We will cover:

- Randall-Sundrum 2 model with delta-function brane.
- Domain walls (kinks/solitons).
- Trapping scalar and fermion fields to a domain wall.
- Using Dvali-Shifman mechanism to trap gauge fields.
- SU(5) grand unified domain-wall model.
- Extending SU(5) to E<sub>6</sub>.
- Cosmology the expansion of a brane universe.

Collaborators: Ray Volkas (Melbourne), Rhys Davies (Oxford), Mark Trodden (U Penn), Kamesh Wali (Syracuse, NY), Aharon Davidson (Ben-Gurion, Israel), Archil Kobakhidze, Gareth Dando (Melbourne).

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Modelling an extra dimension with domain-wall branes

We are inspired by the Randall-Sundrum warped metric solution.

RS1 is a compact extra dimension: provides a solution to the hierarchy problem — lots of work on this model. Branes are string theory like objects. Warped throats, inflation, dark matter, ...

(Randall & Sundrum, PRL83, 3370 (1999))

RS2 is an infinite extra dimension: solves the trapping of low-energy gravity. Not as much interest because it doesn't solve any major problems, just introduces another dimension.

(Randall & Sundrum, PRL83, 4690 (1999))

We will pursue RS2 because it seems a natural extension of 3+1 space. Extra dimensions  $\rightarrow$  extra degrees of freedom to solve problems: weak hierarchy, GUT (proton decay, mass relations), mass hierarchy.

Premise:

- take the standard model and general relativity
- add an *infinite* extra space dimension
- recover the standard model and general relativity at low energies

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$$\mathcal{S} = \int\!\!d^4x \int\!\!dy \sqrt{|g|} \left[ -M^3R + \delta(y) \mathcal{L}_{\rm SM} \right]$$

## RS2 model

Need brane and bulk sources:

$$\mathcal{S} = \int d^4x \int dy \left[ \sqrt{|g|} (-M^3 R - \Lambda_{\text{bulk}}) + \sqrt{|g^{(4)}|} \delta(y) (\mathcal{L}_{\text{SM}} - \Lambda_{\text{brane}}) \right]$$
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(1)

Solve the theory:

- Randall-Sundrum metric ansatz:  $ds^2 = e^{-2k|y|}g^{(4)}_{\mu\nu}dx^{\mu}dx^{\nu} dy^2$
- Solve Einstein's equations ( $\mathcal{L}_{SM} = 0$  and  $R^{(4)} = 0$ ):

$$\Lambda_{\text{bulk}} = -12k^2M^3 \qquad \qquad \Lambda_{\text{brane}} = 12kM^3$$

 $\blacksquare$  Write R in terms of  $R^{(4)} {:}\ R = e^{2k|y|}R^{(4)} - 16k\delta(y) + 20k^2$ 

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• Write R in terms of  $R^{(4)}$ :  $R=e^{2k|y|}R^{(4)}-16k\delta(y)+20k^2$ 

Substitute into (1) and integrate over y:

$$\mathcal{S} = \int \! d^4x \sqrt{|g^{(4)}|} \left[ -\frac{M^3}{k} R^{(4)} + \mathcal{L}_{\mathsf{SM}} \right]$$

A dimensionally reduced theory.

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#### Newton's law

Just need to check Newton's law. Linear tensor fluctuations are:

$$g_{\mu\nu} = e^{-2k|y|}\eta_{\mu\nu} + \sum_{n} h_{n\ \mu\nu}^{(4)}(x^{\mu})\psi_{n}(y)$$

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$$V(r) = -G_N rac{m_1 m_2}{r} \left(1 + rac{\epsilon^2}{r^2}\right)$$
 (where  $\epsilon = 1/k$ )

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- We have what we wanted: a 5D theory that at low energies looks like our 4D universe.
- But almost no new phenomenology.
- Next step: brane forms naturally.



## Thick and smooth RS2

We want to remove the  $\delta(y)$  part of the action:

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Geometry, hence gravity, is 5D. So why not try to make all fields 5D?

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- Then we show how to trap scalars, fermions and gauge fields to the brane.
- Finally we present a 4+1-d SU(5) based extension to the standard model.

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Turn off warped gravity for now — just think about trapping 5D fields.

#### A domain wall as a brane



A *domain-wall* is a (thin) region separating two vacua.

The field  $\phi$  interpolates between two vacua as one moves along the extra-dimension.

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The field  $\phi$  interpolates between two vacua as one moves along the extra-dimension.

Lagrangian for 
$$\phi(x^{\mu}, y)$$
:

$$\mathcal{L} = \frac{1}{2} \partial_M \phi \; \partial^M \phi - V(\phi)$$

A solution is the *kink*:

$$\phi(y) = v \tanh(\sqrt{2\lambda}vy)$$

It is stable!

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extra dimension

#### More complicated domain walls

These examples use two scalar fields to form the wall.



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In a realistic model, symmetries dictate V.

To determine stability, expand  $\phi$  in normal modes about the background:  $\phi(x,t) = \phi_{bg}(x) + \sum_n \xi_n(x) e^{i\omega_n t}$ . Make sure  $\omega_n^2 \ge 0$ 

# Trapping matter fields

# Trapping scalar fields

Aim: to trap a 5D scalar field  $\Xi(x^{\mu}, y)$  to the brane.

A simple quartic coupling works:

$$S = \int d^4x \int dy \left[ \frac{1}{2} \partial^M \phi \ \partial_M \phi - V(\phi) + \frac{1}{2} \partial^M \Xi \ \partial_M \Xi - W(\Xi) - g \phi^2 \Xi^2 \right]$$

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Expand  $\Xi$  in extra dimensional (Kaluza-Klein) modes:

$$\Xi(x^{\mu}, y) = \sum_{n} \xi_n(x^{\mu}) k_n(y)$$

 $\xi_n$  are the 4D fields,  $k_n$  their extra-dimensional profile. The profiles satisfy a Schrödinger equation:

$$\left(-\frac{d^2}{dy^2} + 2g\phi_{\rm bg}^2\right)k_n(y) = E_n^2k_n(y)$$

The energy eigenvalues  $E_n$  are related to the mass of the 4D field  $\xi_n$ .

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# Trapping via a potential well

The effective potential acts like a well.



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To get 4D theory, substitute mode expansion into action and integrate y:

$$\mathcal{S} = \int \! d^4x \left[ \sum_n \left( \frac{1}{2} \partial^\mu \xi_n \partial_\mu \xi_n - m_n^2 \xi_n^2 \right) + (\text{higher order terms}) \right]$$

Orthonormal basis k<sub>n</sub> ⇒ diagonal kinetic and mass terms.
 m<sub>n</sub> can be tuned.

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# Trapping fermions

We can trap a fermion  $\Psi(x^{\mu}, y)$  to the brane with a Yukawa coupling:

$$\mathcal{S} = \int d^4x \int dy \left[ \frac{1}{2} \partial^M \phi \ \partial_M \phi - V(\phi) + \overline{\Psi} i \Gamma^M \partial_M \Psi - h \phi \overline{\Psi} \Psi \right]$$

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Decompose into left- and right-chiral fields and Kaluza-Klein modes:

$$\Psi(x^{\mu}, y) = \sum_{n} \left[ \psi_{Ln}(x^{\mu}) f_{Ln}(y) + \psi_{Rn}(x^{\mu}) f_{Rn}(y) \right]$$

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Schrödinger equation (mode index n suppressed):



## Gravity and matter fields

Brane (domain wall/kink), trapped scalar and fermion. Plus gravity:

$$S = \int d^4x \int dy \sqrt{|g|} \Big[ -M^3 R - \Lambda_{\text{bulk}} + \frac{1}{2} \partial^M \phi \ \partial_M \phi - V(\phi) \\ + \frac{1}{2} \partial^M \Xi \ \partial_M \Xi - W(\Xi) - g \phi^2 \Xi^2 \\ + \overline{\Psi} i \Gamma^M \partial_M \Psi - h \phi \overline{\Psi} \Psi \Big]$$

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Dimensionally reduce by integrating over y:

$$\begin{split} \mathcal{S} &= \int \! d^4 x \sqrt{|g^{(4)}|} \Big[ -M_{4\mathrm{D}}^2 R^{(4)} + \text{(brane dynamics)} \\ &+ \frac{1}{2} \partial^\mu \xi_n \partial_\mu \xi_n - m_n^2 \xi_n^2 - \tau_{mnop} \xi_m \xi_n \xi_o \xi_p - \text{(brane interactions)} \\ &+ \overline{\psi}_{L0} i \gamma^\mu \partial_\mu \psi_{L0} + \overline{\psi}_n (i \gamma^\mu \partial_\mu - \mu_n) \psi_n - \text{(brane interactions)} \Big] \end{split}$$

4D parameters ( $M_{4D}$ ,  $m_n$ ,  $\tau_{mnop}$ ,  $\mu_n$ , brane dynamics) determined by eigenvalue spectra and overlap integrals.

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#### Warped matter

Warped metric  $ds^2 = e^{-2\sigma(y)}\eta_{\mu\nu}dx^{\mu}dx^{\nu} - dy^2$  modifies profile equation:

$$\left(-\frac{d^2}{dy^2} + 5\sigma'\frac{d}{dy} + 2\sigma'' - 6\sigma'^2 + U(y)\right)f_{Ln}(y) = m_n^2 e^{2\sigma}f_{Ln}(y)$$

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Conformal coordinates  $ds^2 = e^{-2\sigma(y(z))}(\eta_{\mu\nu}dx^{\mu}dx^{\nu} - dz^2)$ . Rescale  $f_{Ln}(y) = e^{2\sigma}\tilde{f}_{Ln}(z)$ :

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Matter trapping potentials are warped down.

- Finite bound state lifetimes.
- Resonances.
- Tiny probability of interaction with continuum.



 $(a \sim 1/M^3 \sim 5D$  Newton's constant)

Modelling an extra dimension with domain-wall branes

# Trapping gauge fields

Need to trap gauge fields or e.g. Coulomb potential would be  $V_{\rm Coulomb} \sim 1/r^2.$ 

Not as simple as a Kaluza-Klein mode expansion:

- Photon and gluons *must* remain massless.
- Need to preserve gauge universality at 3+1-d level.

We use the Dvali-Shifman mechanism, Dvali & Shifman, PLB396, 64 (1997).

The following discussion is based on Dvali, Nielsen & Tetradis, PRD77, 085005 (2008).

# Abelian Higgs model

U(1) gauge theory (think 5d photon). Charged Higgs  $\chi$  (gives mass to photon).





extra dimension
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In the bulk:

- U(1) is broken, massive photon  $\sim M_{\chi}$ .
- Higgs vacuum is a superconductor.
- Electric charges are screened.

On the brane:

- $\blacksquare$  U(1) is restored, massless photon.
- Electric field ends on Higgs vacuum.

Charge screening leaks onto the brane!

#### Using a dual superconductor

SU(2) gauge theory (3 "gluons"). Adjoint Higgs  $\chi^a$  (a = 1, 2, 3) (gives mass).





extra dimension

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In the bulk:

- $\blacksquare$   ${\rm SU}(2)$  is restored, in confining regime.
- Large mass gap  $\sim M_{\chi}$  to colourless state.
- QCD-like vacuum is dual superconductor.

On the brane:

- $\blacksquare$  SU(2) broken to U(1), massless photon.
- Electric field repelled from dual superconductor.

For distances much larger than brane width, electric potential  $\sim 1/r.$ 

#### Dvali-Shifman model

Stabilise the domain wall with an extra uncharged scalar field  $\eta$ :

$$S = \int d^4x \int dy \left[ \frac{-1}{4g^2} G^{aMN} G^a_{MN} + \frac{1}{2} \partial^M \eta \partial_M \eta + \frac{1}{2} (D^M \chi^a)^{\dagger} D_M \chi^a - \lambda (\eta^2 - v^2)^2 - \frac{\lambda'}{2} (\chi^a \chi^a + \kappa^2 - v^2 + \eta^2)^2 \right]$$

- $\eta$  has a kink profile.
- If  $\kappa^2 v^2 < 0$ ,  $\chi$  becomes tachyonic near domain wall (where  $\eta \sim 0$ ).
- True vacuum has  $\chi \neq 0$  near domain wall.
- χ breaks symmetry near wall and confines gauge fields.



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extra dimension y

Can add gravity: self consistently solve  $\sigma$  (warped metric profile),  $\eta$ ,  $\chi$ .

The Dvali-Shifman mechanism:

- Works with any non-Abelian SU(N) theory.
- Assumes the SU(N) theory is confining (not proven for 5D).
- Has gauge universality:
  - Charges in the bulk are connected to the brane by a flux tube.
  - Coupling to gauge fields is independent of extra dimensional profile.

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Obvious choice for SU(N) group is SU(5).

# $\mathrm{SU}(5)$ basics

#### Quantum numbers of the standard model

 $\begin{array}{lll} \text{Representations under} & q_L \sim ({\bf 3},{\bf 2})_{1/3} & u_R \sim ({\bf 3},{\bf 1})_{4/3} & d_R \sim ({\bf 3},{\bf 1})_{-2/3} \\ \text{SU}(3) \times \text{SU}(2)_L \times \text{U}(1)_Y & l_L \sim ({\bf 1},{\bf 2})_{-1} & \nu_R \sim ({\bf 1},{\bf 1})_0 & e_R \sim ({\bf 1},{\bf 1})_{-2} \end{array}$ 



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## Putting it all together

## The SU(5) model

Want the standard model on the brane:  $SU(3) \times SU(2)_L \times U(1)_Y$ . Dvali-Shifman needs a *larger* gauge group in the bulk:

 $\mathop{\rm SU}(5)$  is a perfect fit!

Unify the fermions as usual:  $5^*$ , 10. Higgs doublet goes in a  $5^*$ . Want the standard model on the brane:  $SU(3) \times SU(2)_L \times U(1)_Y$ . Dvali-Shifman needs a *larger* gauge group in the bulk:

#### $\mathrm{SU}(5)$ is a perfect fit!

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Summary:

- 4 + 1-dimensional theory all spatial dimensions the same.
- SU(5) local gauge symmetry,  $\mathbb{Z}_2$  discrete symmetry.
- Field content:
  - gauge fields:  $G_{MN} \sim \mathbf{24}$ .
  - **s**calars:  $\eta \sim \mathbf{1}$ ,  $\chi \sim \mathbf{24}$ ,  $\Phi \sim \mathbf{5}^*$ .
  - fermions:  $\Psi_5 \sim \mathbf{5}^*$ ,  $\Psi_{10} \sim \mathbf{10}$ .

The standard model emerges as a low energy approximation.
Ignore gravity for now.

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The action (without gravity)

The theory is described by:

$$\begin{split} \mathcal{S} &= \int \! d^4 x \int \! dy \bigg[ \frac{-1}{4g^2} G^{aMN} G^a_{MN} + \frac{1}{2} \partial^M \eta \partial_M \eta + \text{Tr} \left( (D^M \chi)^{\dagger} (D_M \chi) \right) \\ &+ (D^M \Phi)^{\dagger} (D_M \Phi) + \overline{\Psi}_5 i \Gamma^M D_M \Psi_5 + \overline{\Psi}_{10} i \Gamma^M D_M \Psi_{10} \\ &- h_{5\eta} \overline{\Psi}_5 \Psi_5 \eta - h_{5\chi} \overline{\Psi}_5 \chi^T \Psi_5 \\ &- h_{10\eta} \operatorname{Tr} (\overline{\Psi}_{10} \Psi_{10}) \eta + 2 h_{10\chi} \operatorname{Tr} (\overline{\Psi}_{10} \chi \Psi_{10}) \\ &- h_- \overline{(\Psi_5)^c} \Psi_{10} \Phi - h_+ (\epsilon \overline{(\Psi_{10})^c} \Psi_{10} \Phi^*) + \text{h.c.} \\ &- (c\eta^2 - \mu_{\chi}^2) \operatorname{Tr} (\chi^2) - d\eta \operatorname{Tr} (\chi^3) \\ &- \lambda_1 \left[ \operatorname{Tr} (\chi^2) \right]^2 - \lambda_2 \operatorname{Tr} (\chi^4) - l(\eta^2 - v^2)^2 \\ &- \mu_{\Phi}^2 \Phi^{\dagger} \Phi - \lambda_3 (\Phi^{\dagger} \Phi)^2 - \lambda_4 \Phi^{\dagger} \Phi \eta^2 \\ &- 2\lambda_5 \Phi^{\dagger} \Phi \operatorname{Tr} (\chi^2) - \lambda_6 \Phi^{\dagger} (\chi^T)^2 \Phi - \lambda_7 \Phi^{\dagger} \chi^T \Phi \eta \bigg] \end{split}$$

with kinetic, brane trapping, mass and Dvali-Shifman terms.

#### Split fermions

Let  $\Psi_{nY}$  be the components of  $\Psi_5$  and  $\Psi_{10}$  (n = 5, 10, Y = hypercharge of component), e.g.  $\Psi_5 \supset \Psi_{5,-1} = l_L$ . Dirac equation:

$$\left[i\Gamma^M\partial_M - h_{n\eta}\eta(y) - \sqrt{\frac{3}{5}}\frac{Y}{2}h_{n\chi}\chi_1(y)\right]\Psi_{nY}(x^\mu, y) = 0$$

Each  $\Psi_{nY}$  is a non-chiral 5D field: need to extract the confined left-chiral zero-mode (recall the mode expansion and Schrödinger equation approach):

 $\Psi_{nY}(x^{\mu},y) = \psi_{nY,L}(x^{\mu})f_{nY}(y) + \text{massive modes}$ 

#### Split fermions

Let  $\Psi_{nY}$  be the components of  $\Psi_5$  and  $\Psi_{10}$  (n = 5, 10, Y = hypercharge of component), e.g.  $\Psi_5 \supset \Psi_{5,-1} = l_L$ . Dirac equation:

$$\left[i\Gamma^M\partial_M - h_{n\eta}\eta(y) - \sqrt{\frac{3}{5}}\frac{Y}{2}h_{n\chi}\chi_1(y)\right]\Psi_{nY}(x^\mu, y) = 0$$

Each  $\Psi_{nY}$  is a non-chiral 5D field: need to extract the confined left-chiral zero-mode (recall the mode expansion and Schrödinger equation approach):

$$\Psi_{nY}(x^{\mu}, y) = \psi_{nY,L}(x^{\mu})f_{nY}(y) + \text{massive modes}$$

The effective Schrödinger potential  $\xi$ depends on Y. Thus each component  $\psi_{nY,L}$  has a different profile  $f_{nY}$ .

dimensionless coordinate ky

### Split Higgs

 $\Phi$  contains the Higgs doublet  $\Phi_w$  and a coloured triplet  $\Phi_c$ . Mode expand  $\Phi_{w,c}(x^{\mu}, y) = \phi_{w,c}(x^{\mu})p_{w,c}(y)$ . Schrödinger equation for  $p_{w,c}$  is:

$$\left(-\frac{d^2}{dy^2} + \frac{3Y^2}{20}\lambda_6\chi_1^2 + \sqrt{\frac{3}{5}}\frac{Y}{2}\lambda_7\eta\chi_1 + \dots\right)p_{w,c}(y) = m_{w,c}^2p_{w,c}(y)$$

Critical that ground states have:

- $m_w^2 < 0$  to break electroweak symmetry.
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Critical that ground states have:

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- $m_c^2 > 0$  to preserve QCD.

Large enough parameter space to allow this.



Standard model parameters are computed from overlap integrals.

With one generation of fermions, parameters are easy to fit.

The model overcomes the major SU(5) obstacles:

- $m_e = m_d$  not obtained due to naturally split fermions.
- Coloured Higgs induced proton decay is suppressed.
- Gauge coupling constant running modified due to Kaluza-Klein modes appearing (not analysed yet).

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Adding gravity:

- Solve for warped metric, kink and Dvali-Shifman background.
- Continuum fermion and scalar modes are highly suppressed on the brane.
- Main features remain.

## Going beyond SU(5)

One promising extension is to the  $\mathrm{E}_{6}$  group:

- $E_6 \rightarrow SO(10)$  in the bulk.
- $SO(10) \rightarrow SU(5)$  on the brane due to clash-of-symmetries (CoS) and Dvali-Shifman.
- Can eliminate kink scalar field  $\eta$ .
- Can unify  $\Psi_5$  and  $\Psi_{10}$ .
- Large reduction of free parameters.





## Domain-wall cosmology

#### The scale factor on a brane

Most important cosmological fact: expanding universe.

FRW:  $ds^2 = -dt^2 + a^2(t) d\mathbf{x} \cdot d\mathbf{x}$ 

Energy density  $\rho$  dictates expansion:

$$H = \frac{\dot{a}}{a} \qquad \qquad H^2 = \frac{8\pi G}{3}\rho$$

(Hubble parameter)

(spatially flat universe)



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Η	=	à
		$\overline{a}$

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FRW for a brane:  $ds^2 = -n^2(t, y)dt^2 + a^2(t, y) d\mathbf{x} \cdot d\mathbf{x} + dy^2$ 

$$H_{\rm brane}^2 = \frac{8\pi G}{3} \left(1 + \frac{\rho}{2\sigma}\right)\rho$$

 $\sigma$  is the energy density (tension) of the brane;  $\sigma \gg (1 MeV)^4$  from BBN. (Binétruy, Deffayet & Langlois, Nucl. Phys. B565, 269 (2000))

D.P. George

#### Species-dependent scale factor

$$ds^{2} = -n^{2}(t, y)dt^{2} + \frac{a^{2}(t, y)}{a}d\mathbf{x} \cdot d\mathbf{x} + dy^{2}$$

Each slice at constant y has a different scale factor.



$$ds^{2} = -n^{2}(t, y)dt^{2} + \frac{a^{2}(t, y)}{a}d\mathbf{x} \cdot d\mathbf{x} + dy^{2}$$

Each slice at constant y has a different scale factor.



For thick branes, different species (electron, quark, KK mode) experience different expansion rates.

$$a_{\rm sp}(t) = a_0(t) \left( 1 + \frac{\dot{\rho}}{2\sigma H_0} I_{\rm sp}(t) \right) \label{eq:asp}$$

$$I_{\rm sp}(t) = \frac{\int f_{\rm sp}^2(t,y) \, dy}{\int f_{\rm sp}^2(t,y) \, e^{-\sigma |y|/6M_5^3} dy} - 1$$

Different localisation profile  $f_{sp}$  for different energies! No chiral fermions!

Our universe may be a brane residing in a higher dimensional bulk.

We have investigated the possibility that the brane is in fact a domain-wall.

Main results:

- Scalar field: forms stable domain wall.
- RS2 warped metric: traps gravity.
- Dvali-Shifman mechanism: traps gauge fields.
- $\blacksquare$  SU(5) domain-wall model: overcomes major SU(5) obstacles.
- $E_6$  extension: unify fields and reduce parameters.
- Cosmology: species-dependent expansion rate.

Extra-dimensions a real possibility! Will be tested by LHC (e.g. KK modes) and cosmological observations. Open questions:

- Further analysis of the cosmology.
- Completion of  $E_6$  model.

Incorporate SUSY: (G. R. Dvali & M. A. Shifman, *Nucl. Phys.* **B504** (1997) 127–146)

- SUSY domain wall  $\rightarrow$  dimensional reduction + SUSY breaking.
- SUSY broken only on domain wall, where 4D particles are localised.
- 5D superfield projected to localised 4D fermion (index theorem).
- Solution to DM and hierarchy problems within DW construction.
- Phenomenological predictions at LHC energies.